

Godzilla Geometry aka Cosmic Distance Ladder



Determining Size & Distance Indirectly



Randall H. Landsberg

Director Education & Outreach- KICP Director Public Outreach Dept. Astronomy KICP- University of Chicago









- Context of the Godzilla Geometry Lab
- Empirically Determining the Relationship of Apparent Size to Distance
 - Data Collection
 - Student Data
 - Testing the Model





 Extending the Lesson to Celestial Objects





Space Explorers Science Enrichment Program

- Multi Year Commitment to Inner-City Middle & HS Students
- Weekly Hands-On Laboratories
 - On U. Chicago Campus
 - Graduate Student Instructor
- Enrichment Trips
- Residential Institutes @ Yerkes
 Observatory (WI)
 - Winter Institute 3 Days
 - Summer Institute 1 Week
 - 3 Day Labs, 3 PM Labs, & Presentations
 - Theme 2003 "How Big is the Universe?"
- SMET Success Stats. 500% Better CPS/ National 2007 NSTA Meeting - Landsberg













- Distance Ladder = Pillar of Astronomy
 - **BIGGER** = Closer
 - smaller = Further











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Empirically Measure Apparent Size Vs. Distance

- Objects of Know Size
 Godzilla & Golf Flag
- Know Distances
 10 to 500 feet
- Digital Camera

 Focus Fixed at Infinity
- Student Challenges
 - Indexing Photos



- Mechanics of Measuring Big Distances & Pixels
- Graphing



Data Collection











Graphing & Data Analysis



- •Apparent Size Pictures to Numbers of Pixels
 - Any Program that Counts Pixels
 - Hands On Universe (1st converted jpeg's to fits)





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- Same Godzilla "Unknown" Distance
 - Photo With Same Conditions
 - Mark Location
 - Measure Apparent Size (# Pixles)
 - Determine Distance via Graph
 - Where # Pixels Hit Curve
 - Compare to Marker Distance

























- "Different" Godzilla Known Size
 - Analogous to Astronomy (e.g., different galaxy)
 - Assumes All Godzillas Are Same Size
- Determine Height of Great Refractor Dome
 - Relative Apparent Size
 - # Pixels
 - Vs # Pixels for Know Object Next to it
- Compare to Known Dome Height

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Reinforcing Challenges

- Monster Movie Challenge
 - Make a 3" Godzilla Model Appear as Big as a Person
 - Determine Apparent Size 3" (pixels)
 - Estimate Necessary Separation
 - Test with Camera
- Distance to Another Golf Flag on Golf Course
 - Calculate Distance & Then Measure
 - Assume Similar Objects Same Size
 - Like Rung of Cosmic Distance Ladder
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Climbing the Distance Ladder







- Determine Distance to M31(Andromeda)
 - Assume Globular Clusters Belonging to M31 & the Milky Way are of Comparable Size
- Translate to Absolute Distances
 - Given Exact Distance to Original Globular Cluster
- Determining Distance to Galaxy Clusters
 - Assume Galaxies in Distant Clusters of Galaxies are of Similar Size to M31



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- 1.Assemble Massive M31 Mosaic
- 2. Identify Globular Clusters



- 3.Measure Average Size of Globular Clusters
- 4.Measure the Size of a Nearby Globular Cluster in the Milky Way Galaxy (M3)
- 5.Assume: Small Angle Approximation Valid, SDSS Data Constant Conditions
- 6.Determine Relative Distance (Apparent Sizes)
- 7.Determine Absolute Distance Given Distance to M3/NGC5272 (33,900 Lyr, 10.4kpc)





SDSS Mosaic M31 76 - images hint align by run number and column

HUGE!







SDSS Mosaic M31 76 - images hint align by run number and column

HUGE!





One of 76 M31 Pieces Identify GCs Measure

Globular Cluster M3

Globular cluster M3 - distance is 10,400 parsecs (this comes from RR Lyrae stars). Its angular size is 18 minutes of arc (you can calculate its size from thi

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33,900 Lyr vs 2.53 (0.07) Mly M31

> <u>Quick Dirty</u> - 7px vs 375pixels ~ 54 times Bigger

M3 10.4kpc M31 775kpc ~75 times further









- Break Down of Small Angel Approximation
 - Too BIG, Too Close (typically not a problem in astronomy - huge distances)
- Lens Distortions (not flat)
 - take photo to demo (eg brick wall)
- Data Needs to be Taken w/Same Conditions (Apples to Apples)
 - Camera Must Be Set Same Way (zoom to infinity)
 - Homogenous Survey Data Set Like SDSS



Summing Up

NSF

- Visceral Feel for Abstract Concept
 - Demonstrates Indirect Determinations
- Areas for Extensions
 - Angular Scale
 - Small Angle Approximation
 - Resolution
- Real Data Available
 SDSS Skyserver
- Documented on Web
 - http://kicp.uchicago.edu/nsta







- Kavli Institute of Cosmological Physics
 - National Science Foundation (NSF)
 - NSF PHY-0114422
 - Kavli Foundation
- Data



- Sloan Digital Sky Survey (sdss.org)
- Office of Special Programs (UC)

- Partners Space Explorers Program

- Mark SubbaRao (UC/Adler)
 - Co-taught

Online resources http:kicp.uchicago.edu/nsta

The End



Angular Size & Small Angles

- Favorite of Astronomers
- Do Not Need Distance
- Independent of Telescope
- (far far away limit of small angle approximation)



Determining Pixels/Degree

- Angular Scale φ
 - Photo Know Height & Know Distance
 - tan (ϕ) = opposite/adjacent= height/distance
 - ϕ (radians) = arctan (opposite/adjacent)
 - ϕ (degrees)= ϕ (radians)x180/ π
 - ϕ (degrees)= arctan (opposite/adjacent) x180/ π
- Pixels
 - Count the number of vertical pixels image takes up
- Pixels/Degree
 - (angular size ϕ) = (number of pixels)

Nikon CoolPix 4500

38 pixels/ degree



Pictures to Pixels to Apparent Angles



"4" Mega Pixel

 CCD Chip 2,272 wide x 1,704 tall
 2,272x1,704 = 3,871,488 pixels